Journal of The Tensor Society (J.T.S.) ISSN: 0974-5424 Vol. 17 (2023), page 54-59

# On the Lanczos Potential

Ravi Panchal $^{1*}$ , A.H. Hasmani $^{1**}$  and J. López-Bonilla $^{2+},$ 

<sup>1</sup>Department of Mathematics, Sardar Patel University, Vallabh Vidyanagar-388120, India. <sup>2</sup>ESIME-Zacatenco, Instituto Politécnico Nacional, Edif. 4, 1er. Piso, Col.Lindavista CP 07738, CDMX, México E-mail: <sup>∗</sup> ravipanchal1712@spuvvn.edu, ∗∗ah hasmani@spuvvn.edu, <sup>+</sup>jlopezb@ipn.mx

### Abstract

For arbitrary geometries of Petrov types  $III$ ,  $N$  and  $O$  we construct the Lanczos potential for the corresponding Weyl tensor. This will provide a technique to construct Lanczos potential in non-vacuum cases.

Keywords: Conformal tensor, Lanczos generator, Canonical null tetrad, Petrov classification, Newman-Penrose formalism, 2-spinors.

## 1 Introduction

The Lanczos potential  $K_{\mu\nu\alpha}$  [1, 2, 3, 4, 5, 6, 7, 8] satisfies the algebraic symmetries:

$$
K_{\mu\nu\alpha} = -K_{\nu\mu\alpha}, \quad K_{\mu\nu\alpha} + K_{\nu\alpha\mu} + K_{\alpha\mu\nu} = 0, \quad K^{\mu\nu}{}_{\nu} = 0 \tag{1.1}
$$

and it generates the Weyl tensor [9, 10] via the expression [11]:

$$
-C_{\mu\nu\alpha\beta} = K_{\mu\nu\alpha;\beta} - K_{\mu\nu\beta;\alpha} + K_{\alpha\beta\mu;\nu} - K_{\alpha\beta\nu;\mu} + \frac{1}{2} [(K_{\mu\beta} + K_{\beta\mu})g_{\nu\alpha} + (K_{\nu\alpha} + K_{\alpha\nu})g_{\mu\beta} - (K_{\mu\alpha} + K_{\alpha\mu})g_{\nu\beta} - (K_{\nu\beta} + K_{\beta\nu})g_{\mu\alpha}] + \frac{2}{3} K^{\lambda\sigma}{}_{\lambda;\sigma}, \quad \text{where } K_{\mu\nu} = K_{\mu\sigma\nu}{}^{;\sigma} - K_{\mu\sigma}{}^{\sigma}{}_{;\nu}
$$
\n(1.2)

Though the introduction of such a tensor provides an analogy of gravitation and electromagnetism, the possible physical meaning of  $K_{\mu\nu\alpha}$  in general relativity, is an open problem. Here we exhibit the structure of the Lanczos generator for arbitrary spacetimes of Petrov types  $O, N$  and  $III$  [12], via the Newman-Penrose technique [9, 10, 13, 14] and tensor and spinor formalisms [15]. Given a geometry we must construct the corresponding Lanczos potential, which is difficult in the general case except for certain Petrov types, for example, if in (1.2) we use the expression:

$$
K_{\mu\nu\alpha} = N_{\mu}g_{\nu\alpha} - N_{\nu}g_{\mu\alpha},\tag{1.3}
$$

in terms of the metric tensor with  $N_{\mu}$  arbitrary, we obtain  $C_{\mu\nu\alpha\beta} = 0$ , that is, (1.3) is a Lanczos generator for any spacetime type O.

This paper is dedicated to late Prof. Oscar Chavoya Aceves (1957-2023) for his work in the field of Relativity theory.

### 2 Algebraic symmetries of the Lanczos potential

If  $a_{\mu\nu\alpha}$  is an arbitrary tensor, then the following tensor [16]:

$$
\tilde{K}_{\mu\nu\alpha} = a_{\mu\nu\alpha} - a_{\nu\mu\alpha} + a_{\alpha\nu\mu} - a_{\alpha\mu\nu} \tag{2.1}
$$

satisfies the symmetries:

$$
\tilde{K}_{\mu\nu\alpha} = -\tilde{K}_{\nu\mu\alpha}, \quad \tilde{K}_{\mu\nu\alpha} + \tilde{K}_{\nu\alpha\mu} + \tilde{K}_{\alpha\mu\nu} = 0,
$$
\n(2.2)

where we may employ:

$$
a_{\mu\nu\alpha} = \frac{1}{3} F_{\mu\nu;\alpha}, \quad F_{\mu\nu} = -F_{\nu\mu}, \tag{2.3}
$$

whose application in (4) gives the expression:

$$
\tilde{K}_{\mu\nu\alpha} = \frac{1}{3} (2F_{\mu\nu;\alpha} + F_{\alpha\nu;\mu} - F_{\alpha\mu;\nu}),\tag{2.4}
$$

and if now we ask the condition  $\tilde{K}_{\mu\nu\alpha} = 0$ , then from (2.4) we obtain a tensor verifying the algebraic symmetries (1.1) of the Lanczos potential [11]:

$$
K^{\mu\nu\alpha} = \frac{1}{3} (2F^{\mu\nu;\alpha} + F^{\alpha\nu;\mu} - F^{\alpha\mu;\nu} + F^{\nu\lambda}_{;\lambda} g^{\alpha\mu} - F^{\mu\lambda}_{;\lambda} g^{\alpha\nu}), \tag{2.5}
$$

which we shall use in our study.

If in (4) we employ the option  $a_{\mu\nu\alpha} = F_{\mu\nu} k_\alpha$ ,  $F_{\mu\nu} = -F_{\nu\mu}$  with the properties:

$$
F_{\mu\nu}k_{\alpha} + F_{\nu\alpha}k_{\mu} + F_{\alpha\mu}k_{\nu} = 0, \quad \text{where } F_{\mu\nu}k^{\nu} = 0,
$$
\n(2.6)

and the trace is eliminated, then we obtain that  $K_{\mu\nu\alpha} \propto F_{\mu\nu} k_{\alpha}$ , which allows construction of the Lanczos generator for plane gravitational waves [spacetime of Petrov type N] [17].

# 3 Lanczos potential for arbitrary 4-geometries with Petrov types  $O, N$  and  $III$

If we select (2.5) for arbitrary  $F_{\mu\nu} = -F_{\nu\mu}$ , then (1.2) implies the relation:

$$
C_{\mu\nu\alpha\beta} = C_{\sigma\nu\alpha\beta} F^{\sigma}_{\ \mu} - C_{\sigma\mu\alpha\beta} F^{\sigma}_{\ \nu} + C_{\sigma\beta\mu\nu} F^{\sigma}_{\ \alpha} - C_{\sigma\alpha\mu\nu} F^{\sigma}_{\ \beta} \tag{3.1}
$$

which identically vanishes for any conformally flat space, that is,  $(2.5)$  is a Lanczos potential for any spacetime with Petrov type O.

Now we consider two Petrov types in the canonical null tetrad [9, 13, 18], for certain  $F_{\mu\nu}$ :

(a) Type  $N$ :

$$
C_{\mu\nu\alpha\beta} = \psi_4 V_{\mu\nu} V_{\alpha\beta} + \bar{\psi}_4 \bar{V}_{\mu\nu} \bar{V}_{\alpha\beta},
$$
\n(3.2)

where  $V_{\mu\nu} = l_{\mu}m_{\nu} - l_{\nu}m_{\mu}$  and

$$
F_{\alpha\beta} = q(n_{\alpha}l_{\beta} - n_{\beta}l_{\alpha}).
$$
\n(3.3)

Thus, (3.1) takes the form  $C_{\mu\nu\alpha\beta} = 2qC_{\mu\nu\alpha\beta}$ , therefore (2.5) is a Lanczos potential with (3.3) for  $q = 1/2$ .

(b) Type  $III$ :

$$
C_{\mu\nu\alpha\beta} = \psi_3 (V_{\mu\nu} M_{\alpha\beta} + M_{\mu\nu} V_{\alpha\beta}) + \bar{\psi}_3 (\bar{V}_{\mu\nu} \bar{M}_{\alpha\beta} + \bar{M}_{\mu\nu} \bar{V}_{\alpha\beta}), \tag{3.4}
$$

where  $M_{\mu\nu} = m_{\mu} \bar{m}_{\nu} - m_{\nu} \bar{m}_{\mu} + n_{\mu} l_{\nu} - n_{\nu} l_{\mu}$ . Then (3.1), (3.3) and (3.4) give the relation  $C_{\mu\nu\alpha\beta} = qC_{\mu\nu\alpha\beta}$ , that is, (2.5) is a Lanczos generator with (3.3) for  $q=1$ .

Hence, the Lanczos potential for arbitrary Petrov types  $N$  and  $III$  spacetimes has the structure (2.5) if we employ the corresponding canonical null tetrad and  $F_{\mu\nu}$  is given by (3.3) with  $q = 1/2$  and  $q = 1$ , respectively; in Petrov type O geometries, we can use (2.5) with any  $F_{\alpha\beta}$ . The construction of  $K_{\mu\nu\alpha}$  for arbitrary 4-spaces of types I, II, and D, is an open problem.

## 4 About Lanczos Spinor

From (2.5):

$$
S_{\mu\nu\alpha} = K_{\mu\nu\alpha} + i^* K_{\mu\nu\alpha} = \frac{1}{3} (2S_{\mu\nu;\alpha} + S_{\alpha\nu;\mu} - S_{\alpha\mu;\nu} + S_{\nu\lambda}^{i\lambda} g_{\alpha\mu} - S_{\mu\lambda}^{i\lambda} g_{\alpha\nu}), \tag{4.1}
$$

such that  $S_{\mu\nu} = F_{\mu\nu} + i * F_{\mu\nu}$ , with the participation of the dual tensor:

$$
{}^*K_{\mu\nu\alpha} = \frac{1}{2} \eta_{\mu\nu\lambda\beta} K^{\lambda\beta}_{\ \alpha} \quad \text{and} \quad {}^*F_{\mu\nu} = \frac{1}{2} \eta_{\mu\nu\lambda\beta} F^{\lambda\beta} \tag{4.2}
$$

From  $(4.1)$  the corresponding Lanczos spinor  $[19, 20, 21, 22, 23]$  is given by:

$$
3L(ABC\dot{D}) = \nabla_{A\dot{D}} \varphi_{BC} + \nabla_{B\dot{D}} \varphi_{CA} + \nabla_{C\dot{D}} \varphi_{AB},\tag{4.3}
$$

which implies the following equations in the Newman-Penrose (NP) formalism [9, 13, 14, 21, 24]:

$$
\Omega_0 = D\phi_0 + 2(-\epsilon\phi_0 + \kappa\phi_1), \quad 3\Omega_1 = \bar{\delta}\phi_0 + 2[D\phi_1 - (\alpha + \pi)\phi_0 + \rho\phi_1 + \kappa\phi_2], \n\Omega_3 = \bar{\delta}\phi_2 + 2(-\lambda\phi_1 + \alpha\phi_2), \quad 3\Omega_2 = D\phi_2 + 2[\bar{\delta}\phi_1 - \lambda\phi_0 - \pi\phi_1 + (\rho + \epsilon)\phi_2], \n\Omega_4 = \delta\phi_0 + 2(-\beta\phi_0 + \sigma\phi_1), \quad 3\Omega_5 = \Delta\phi_0 + 2[\delta\phi_1 - (\gamma + \mu)\phi_0 + \tau\phi_1 + \sigma\phi_2], \n\Omega_7 = \Delta\phi_2 + 2(-\nu\phi_1 + \gamma\phi_2), \quad 3\Omega_6 = \delta\phi_2 + 2[\Delta\phi_1 - \nu\phi_0 - \mu\phi_1 + (\beta + \tau)\phi_2],
$$
\n(4.4)

for the NP components of Lanczos potential in terms of the spin coefficients and the NP projections of  $F_{\mu\nu}$ .

The work [25] used the canonical null tetrad [9, 18, 26] to determine the NP components of  $K_{\mu\nu\alpha}$ , that is, a solution of the Weyl-Lanczos equations [8, 20, 21, 25, 27] thus:

$$
\Omega_0 = qk, \qquad \Omega_3 = -q\lambda, \qquad \Omega_4 = q\sigma, \qquad \Omega_7 = -q\nu
$$
  
\n
$$
\Omega_1 = \frac{q}{3}\rho, \qquad \Omega_2 = -\frac{q}{3}\pi, \qquad \Omega_5 = \frac{q}{3}\tau, \qquad \Omega_6 = -\frac{q}{3}\mu \qquad (4.5)
$$

for arbitrary spacetimes with Petrov types III and N for  $q = 1$  and  $q = 1/2$ . respectively. It is easy to see that the relations (4.4) imply (4.5) if  $\phi_0 = \phi_2 = 0$  and  $\phi_1 = \frac{q}{2}$  $\frac{q}{2}$ , hence  $F_{\mu\nu}$  has the structure (3.3) and  $S_{\mu\nu} = qM_{\mu\nu}$ .

## 5 Conclusion

Due to non-linearity in Weyl-Lanczos relations, their solution is not unique and it becomes difficult to solve them for spacetimes of general nature. In this work we have derived Lanczos potential for metric of Petrov types III, N and O. It is pending to determine the Lanczos potential for Petrov types  $I, II$  and  $D$  in the general case, however, for these spacetimes  $K_{\mu\nu\alpha}$  has been constructed for specific metrics [28, 29, 4, 5, 6, 8, 30, 31, 32, 33, 34, 35] and in almost all of these situations the NP components  $\Omega_r$  have turned out to be linear combinations of the spin coefficients if we properly choose the corresponding null tetrad. Equation (2.5) may be seen as interaction between gravitational field in vacuum and electromagnetic field, which requires further investigation. There are various proposals [30, 31, 36, 37, 38] for the possible physical meaning of the Lanczos spintensor, which we consider important because  $K_{\mu\nu\alpha}$  is present in all Riemannian 4-geometries and therefore it is necessary to know its role in the description of the gravitational field.

### Acknowledgments

RP is thankful to Sardar Patel University for providing financial support under SEED Money Grant.

### References

- [1] C. Lanczos. The Splitting of the Riemann Tensor. Rev. Mod. Phys., 34:379–389, 1962.
- [2] C. Lanczos. The Riemannian tensor in four dimensions. Annales de la Faculté des Sciences de Universite de Clermont-Ferrand, 8:167–170, 1962.
- [3] Zoltán Perjés. The works of Kornél Lánczos on the Theory of Relativity. In  $A$ Panorama of Hungarian Mathematics in the Twentieth Century I, pages 415– 425. Springer, 2006.
- [4] R.G. Vishwakarma. A new Weyl-like tensor of geometric origin. Journal of Mathematical Physics, 59(4), 2018.
- [5] I. Guerrero-Moreno, J. López-Bonilla, R. López-Vázquez, and S. Vidal-Beltrán. Lanczos generator in terms of the conformal tensor in Gödel and type D vacuum geometries. American-Eurasian J. of Sci. Res., 13(4):67–70, 2018.
- [6] J. López-Bonilla, R. López-Vázquez, and S. Vidal-Beltrán. Bach, Cotton and Lanczos tensors in Gödel geometry. African J. Basic and Appl. Sci.,  $11(1):18$ . 2019.
- [7] J. López-Bonilla, J. Morales, and G. Ovando. A comment on the Lanczos conformal Lagrangian. World Eng. Appl. Sci. J., 10(2):61, 2019.
- [8] Zafar Ahsan. The Potential of Fields in Einstein's Theory of Gravitation. Springer, 2019.
- [9] Hans Stephani, Dietrich Kramer, Malcolm MacCallum, Cornelius Hoenselaers, and Eduard Herlt. Exact Solutions of Einstein's Field Equations. Cambridge University Press, 2009.
- [10] Pankaj Sharan. Spacetime, geometry and gravitation. Springer, 2009.
- [11] M. Novello, E. Bittencourt, and J.M. Salim. The Quasi-Maxwellian equations of general relativity: applications to perturbation theory. Brazilian Journal of Physics, 44(6):832–894, 2014.
- [12] M. Acevedo, M. Enciso-Aguilar, and J. L´opez-Bonilla. Petrov classification of the conformal tensor. Electronic Journal of Theoretical Physics, 3(9):79–82, 2006.
- [13] Ezra Newman and Roger Penrose. An Approach to Gravitational Radiation by a Method of Spin Coefficients. Journal of Mathematical Physics, 3(3):566–578, 1962.
- [14] A.H. Hasmani. Algebraic Computation of Newmann-Penrose Scalars in General Relativity using Mathematica. Journal of Science, 1:82–83, 2010.
- [15] K.P. Tod. Spinors and Spin Coefficients. Encyclopedia of Mathematical Physics. IV, 2006.
- [16] David Lovelock and Hanno Rund. Tensors, differential forms, and variational principles. Courier Corporation, 1989.
- [17] J.L. L´opez-Bonilla, G. Ovando, and J.J. Pena. A Lanczos potential for plane gravitational waves. Foundations of Physics Letters, 12:401–405, 1999.
- [18] J López-Bonilla, R López-Vázquez, J Morales, and G Ovando. Petrov types and their canonical tetrads. Prespacetime Journal, 7(8):1176–1186, 2016.
- [19] W. F. Maher and J. D. Zund. A Spinor Approach to the Lanczos Spin Tensor. Il Nuovo Cimento A (1965-1970), 57(4):638–648, 1968.
- [20] J.D. Zund. The theory of the Lanczos spinor. Annali di Matematica Pura ed Applicata, 104:239–268, 1975.
- [21] Peter O'Donnell. Introduction to 2-Spinors in General Relativity. World Scientific, 2003.
- [22] J. López-Bonilla, R. López-Vázquez, J. Morales, and G. Ovando. Maxwell, lanczos and weyl spinors. Prespacetime Journal, 6(6):509–520, 2015.
- [23] J. Morales, G. Ovando, J. López-Bonilla, and R. López-Vázquez. Lanczos spinor. African J. of Basic and Appl. Sci., 11(4):87–90, 2019.
- [24] J. López-Bonilla, R. López-Vázquez, J. Morales, and G. Ovando. Spin coefficients formalism. Prespacetime Journal, 6(8):697–709, 2015.
- [25] Gonzálo Ares de Parga, Oscar Chavoya A., and José L. López Bonilla. Lanczos Potential. Journal of Mathematical Physics, 30(6):1294–1295, 1989.
- [26] Jerzy Plebanski and Andrzej Krasinski. An introduction to general relativity and cosmology. Cambridge University Press, 2006.
- [27] Ahmet Baykal and Burak Unal. A derivation of Weyl-Lanczos equations. The European Physical Journal Plus, 133:1–14, 2018.
- [28] Zafar Ahsan and Mohd Bilal. On the Lanczos Potential for Petrov Type III Spacetimes. Journal of Tensor Society, 6(2):127–134, 2012.
- [29] Zafar Ahsan, Mohd Bilal, and Musavvir Ali. Petrov Type D Spacetimes and Lanczos Potential. Journal of the Tensor Society, 11(01):97–105, 2017.
- [30] V. Gaftoi, G. Ovando, and J.L. Lopez-Bonilla. Lanczos potential for a rotating black hole. Nuovo Cimento della Societa Italiana di Fisica [Sezione] B, 113:1493–1496, 1998.
- [31] J. López-Bonilla and G. Ovando. Lanczos Spintensor for the Godel Metric. General Relativity and Gravitation, 31(7):1071–1074, 1999.
- [32] A. H. Hasmani and Ravi Panchal. Lanczos Potential for Some Non-Vacuum Spacetimes. The European Physical Journal Plus, 131(9):1–6, 2016.
- [33] A.H. Hasmani, A.C. Patel, and Ravi Panchal. Lanczos potential for Weyl metric. PRAJNA-E-Journal of Pure and Applied Sciences, Accepted, 24:11–14, 2017.
- [34] Ravi Panchal, A.C. Patel, and A.H. Hasmani. On Lanczos Potential for Spherically Symmetric Spacetimes. 13:31–43, 2019.
- [35] J. López-Bonilla, M. Shadab, and Saibal Ray. Factorization of the metric tensor in Kerr geometry. Scientific Voyage, 2(2):19–25, 2021.
- [36] M. D. Roberts. The Physical Interpretation of the Lanczos Tensor. Il Nuovo Cimento B (1971-1996), 110(10):1165–1176, 1995.
- [37] Ram Gopal Vishwakarma. Relativistic potential of Weyl: a gateway to quantum physics in the presence of gravitation? Classical and Quantum Gravity, 37(6):065020, 2020.
- [38] Ram Gopal Vishwakarma. Lanczos potential of Weyl field: interpretations and applications. The European Physical Journal C, 81(2):194, 2021.